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RADIAL VELOCITIES OF 30 STARS

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During the course of several years' work at this observatory a considerable number of miscellaneous measures of radial velocity have accumulated. Some of these are of spectroscopic binaries already announced, while others are of stars for which no previous measures were known. In the case of the latter, a few have been found to be variable in velocity and are so indicated in the column of remarks of the following table or in the text. Some of the binaries could with profit be continued, while others are either too uncertain for measurement or have such a small range that greater dispersion is necessary to successfully tackle them. The stars with their approximate positions for 1900, the number of plates secured, and other data are given in the following table.

Star	$\alpha$		$\delta$		Plates	Remarks
	h	m	°	'		
$\eta$ Piscium.....	1	26	+14	50	40	uncertain binary
$\sigma$ Ceti.....	2	14	- 3	26	4	
35 Arietis.....	2	38	+27	17	5	constant velocity
Boss 678.....	2	54	+51	57	5	suspected binary
Boss 744.....	3	12	+49	43	4	
Boss 783.....	3	22	+49	30	2	
Boss 839.....	3	36	+33	39	7	new spec. binary
42 Persei.....	3	43	+32	48	4	
Boss 898.....	3	49	+47	35	2	
Boss 947.....	4	01	+47	27	5	new spec. binary
$\psi$ Orionis.....	5	22	+ 3	01	16	
$\iota$ Orionis.....	5	30	- 5	59	30	
30 Canis Majoris.....	7	15	-24	47	3	
Boss 2381.....	8	47	- 6	49	2	
36 Lyncis.....	9	07	+43	38	5	
$\zeta$ Leonis.....	10	11	+23	55	3	



Star	$\alpha$		$\delta$		Plates	Remarks
	h	m	°	'		
$\pi$ 8 Virginis.....	11	56	+ 7	10	15	
16 Comæ.....	12	22	+27	22	4	
12 Canum Venaticorum.....	12	51	+38	52	4	new spec. binary
$\tau$ Virginis.....	13	57	+ 2	02	12	constant velocity, - 2
33 Boötis.....	14	35	+44	50	2	
10 Serpentis.....	15	24	+ 2	12	4	
12 Coronæ Borealis.....	15	52	+38	14	3	
21 Ophiuchi.....	16	46	+ 1	23	4	
101 Herculis.....	18	05	+20	01	4	
Boss 4669.....	18	22	+29	46	2	
50 Draconis.....	18	50	+75	20	28	additional measures
Boss 5070.....	19	47	+40	20	7	
13 Vulpeculæ.....	19	49	+23	50	6	
Boss 5535.....	21	28	+60	01	4	

 $\eta$  PISCIIUM

(1900,  $\alpha = 1^{\text{h}} 26^{\text{m}}.1$ ,  $\delta = + 14^{\circ} 50'$ , mag. 3.72, type G5)

In the *Astrophysical Journal*, volume XIX, page 249 and volume XXI, page 313, are given the measures of 15 plates taken by Lord at the Emerson McMillan Observatory in the years 1901 to 1905. The range shown is from +9.5 to +24.9, and Lord suspected that the star was a spectroscopic binary of long period. Giving his plates equal weight would bring the mean velocity about +16.4 km. per second.

From 1897 to 1904 there were 7 plates of the star secured at the Lick Observatory with practically no range in velocity shown. Campbell used the mean +15.5 km. per sec. as the velocity of the star, though he stated that Lord's contention of its spectroscopic character was neither proved nor disproved by his plates.

From 1904 to 1907 Küstner secured 4 plates with a single-prism spectrograph showing no appreciable range and giving a mean velocity of +14.8 km. per second.

In 1906, 1907 and 1908 there were 44 plates secured here as given in the table following. For reasons given in the column of remarks the numbers 606, 624, 1057 and 1254 have not been considered in the discussion. Of the remaining 40 plates, 16 were made with the three-prism universal spectroscope adapted for radial velocity work, 5 with the single-prism long-focus camera, and the remaining 19 with the regular three-prism long-focus camera arrangement. From our early plates it was felt that the variation was real and the period short, and a number of plates were made and measured—the labourious method of applying the Hartman-Cornu formula for each line being used—before it was suspected that some systematic error in the instrument might be the cause of the variations measured. This suspicion is probably the correct view for the universal and single-prism instruments, as in the early stages of the work they were not so perfect as later experience made them.

The range shown by the plates of the regular three-prism instrument is 11 km. but if three of these are omitted, in which only about half the usual number of lines are measured, the greatest range is about 3 km. It would appear only reasonable, then, to assume that if the star is a spectroscopic binary its range of variation is very small. On the assumption that its velocity is constant, the mean of either the universal plates or the single-prism plates is +12.3 km. per sec., while for the three-prism it is +14.1 km. per sec. Weighting the single, universal and three-prism arrangements as 1, 2 and 3 the mean velocity is +13.3 km. per second.

MEASURES OF  $\eta$  PISCUM

Plate	Inst.	Date	Vel.	<i>n</i>	Remarks
403.....	U	1906 Sept. 27-847	+10.7	7	
554.....	"	1907 Jan. 18-543	+ 3.6	26	
562.....	"	" 21-550	+16.3	25	
568.....	"	" 22-604	+10.7	28	
584.....	"	" 28-526	+10.0	28	
588.....	"	" 30-552	+ 5.1	23	
595.....	"	Feb. 4-542	+16.3	17	
598.....	"	" 6-521	+ 6.5	29	
606.....	"	" 7-529	+26.4	10	plate underexposed
608.....	"	" 8-542	+17.9	14	
616.....	"	" 21-487	+ 5.2	18	
617.....	"	" 21-529	+18.1	23	
624.....	"	" 22-500	+ 3.8	20	temperature change
625.....	"	" 25-539	+21.4	24	
634.....	"	" 25-543	+18.6	24	
635.....	"	" 25-580	+18.5	8	
641.....	"	" 27-507	+10.9	25	
642.....	"	" 27-544	+16.0	15	
955.....	I L	July 18-799	+ 0.6	7	
1003.....	III L	Aug. 10-494	+18.4	9	
1036.....	I L	Sept. 6-750	+ 8.0	17	
1041.....	"	" 12-767	+12.9	17	
1051.....	"	" 18-660	+19.1	18	
1054.....	"	" 18-781	+13.2	19	
1057.....	"	" 18-885	+ 8.5	10	plate underexposed
1095.....	III L	Oct. 8-734	+12.2	26	
1096.....	"	" 8-799	+12.2	25	
1102.....	"	" 18-670	+13.0	16	
1103.....	"	" 22-647	+15.1	14	
1104.....	"	" 23-644	+11.8	25	
1105.....	"	" 23-697	+13.9	18	
1107.....	"	" 25-633	+13.2	19	
1164.....	"	Nov. 29-543	+14.3	17	
1165.....	"	" 29-589	+12.3	11	
1173.....	"	Dec. 4-577	+11.4	18	
1187.....	"	" 21-483	+14.1	18	

MEASURES OF  $\eta$  PISCUM—*Concluded*

Plate	Inst.	Date	Vel.	$n$	Remarks
1908					
1205.....	III L	Jan. 1·544	+13·9	16	
1211.....	"	" 3·501	+10·9	10	
1254.....	"	" 22·516	+ 8·5	9	underexposed
1796.....	"	Aug. 19·840	+14·3	13	
1869.....	"	Sept. 7·716	+12·8	21	
1982.....	"	Nov. 21·688	+13·9	12	
2002.....	"	Dec. 5·473	+22·4	8	
2069.....	"	" 23·538	+20·9	10	

 $\sigma$  CETI(1900,  $\alpha = 2^{\text{h}} 14^{\text{m}} \cdot 3$ ,  $\delta = -3^{\circ} 26'$ , mag. var., type Md)

Four spectra of this well known variable star have been made since the extended series at the December 1906 maximum. Apart from the emission band at  $H_{\beta}$ , which was present at the 1906 maximum and is not seen at the time of these observations, the other emission and absorption bands seem to be similar to those of the earlier date. The measures of the sharp  $H_{\gamma}$  and  $H_{\delta}$  emission are fairly reliable, those for the absorption lines only approximate.  $H_{\beta}$  emission seems to appear only at the brighter maxima.

Plate	Date, G.M.T.	Absorption		Emission	
		Vel.	Lines	Vel.	Lines
6786.....	1915, Feb. 17·525	+50	3	+48·7	2
6834.....	Mar. 4·500	+58	4	+45·1	1
8655.....	1918, Sept. 24·703			+49·2	2
8659.....	" 30·708	+70	3	+45·3	2

MEASURES OF  $\alpha$  CETI

$\lambda$	6786		6834		8655		8659		Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.						
4535.965	+67.5	$\frac{1}{2}$	+83.2	$\frac{1}{2}$	.....	.....	.....	.....	.....	.....	.....	.....	.....	.....
4404.927	.....	.....	87.8	$\frac{1}{2}$	.....	.....	+57.8	$\frac{1}{2}$	.....	.....	.....	.....	.....	.....
4395.286	.....	.....	.....	.....	.....	.....	54.1	$\frac{1}{4}$	.....	.....	.....	.....	.....	.....
4325.939	88.5	$\frac{1}{2}$	73.0	$\frac{1}{4}$	.....	.....	67.9	$\frac{1}{4}$	.....	.....	.....	.....	.....	.....
4308.081	69.1	$\frac{1}{4}$	61.0	$\frac{1}{4}$	.....	.....	.....	.....	.....	.....	.....	.....	.....	.....
4340.634 em	78.2	1	70.9	1	+36.4	1	37.5	1	.....	.....	.....	.....	.....	.....
4101.890 em	+71.6	1	+63.2	1	+35.3	1	+31.5	1	.....	.....	.....	.....	.....	.....
Weighted mean	+ 76.24		+ 79.67		.....		+ 59.40		.....	.....	.....	.....	.....	.....
$V_s$	- 25.74		- 21.39		+ 13.45		+ 10.81		.....	.....	.....	.....	.....	.....
$V_z$	- .22		- .25		+ .17		+ .14		.....	.....	.....	.....	.....	.....
Curv.	- .28		- .28		- .28		- .28		.....	.....	.....	.....	.....	.....
Radial Velocity	+ 50.0		+ 57.8		.....		+ 70.1		.....	.....	.....	.....	.....	.....
Emission	+ 48.7		+ 45.1		+ 49.2		+ 45.3		.....	.....	.....	.....	.....	.....

35 ARIETIS

(1900,  $\alpha = 2^h 37^m \cdot 6$ ,  $\delta = +27^\circ 17'$ , mag. 4.58, type B8)

No measures seem to have been published for this star. The lines of hydrogen and helium while fairly broad should give reasonable accuracy in the measures. Plate 3667 is underexposed, so that a variation in velocity cannot be said to be established.

Plate	Date, G.M.T.	Velocity	Lines
2708.....	1909, Aug. 2.837	+17	8
2858.....	Oct. 6.787	+23	7
3650.....	1910, Sept. 14.839	+14	8
3667.....	" 16.758	+35	3
8717.....	1919, Feb. 24.526	+19	7

## MEASURES OF 35 ARIETIS

$\lambda$	2708		2858		3650		3667		8717					
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4861·527	.....		+ 0·3	$\frac{1}{4}$	+ 1·0	$\frac{1}{4}$	.....		+83·6	$\frac{1}{4}$	.....		.....	
4481·400	- 9·3	$\frac{1}{4}$	.....		.....		.....		43·9	$\frac{1}{4}$	.....		.....	
4471·676	-17·6	$\frac{1}{4}$	+ 5·4	$\frac{1}{4}$	- 4·5	$\frac{1}{2}$	- 7·1	$\frac{1}{4}$	38·1	$\frac{1}{2}$	.....		.....	
4388·100	+ 6·7	$\frac{1}{4}$	.....		+ 6·7	$\frac{1}{4}$	.....		44·7	$\frac{1}{4}$	.....		.....	
4340·634	-21·2	$\frac{1}{2}$	+33·2	$\frac{1}{2}$	-20·0	$\frac{1}{2}$	+13·8	$\frac{1}{4}$	48·7	$\frac{1}{2}$	.....		.....	
4143·928	+ 7·2	$\frac{1}{4}$	+21·0	$\frac{1}{4}$	+20·0	$\frac{1}{4}$	.....		.....		.....		.....	
4101·890	- 6·7	$\frac{1}{4}$	-24·9	$\frac{1}{4}$	-13·7	$\frac{1}{4}$	+30·1	$\frac{1}{4}$	58·9	$\frac{1}{4}$	.....		.....	
4026·352	-19·3	$\frac{1}{2}$	0·0	$\frac{1}{2}$	-21·6	$\frac{1}{2}$	.....		+20·0	$\frac{1}{4}$	.....		.....	
3933·825	-17·6	$\frac{1}{4}$	+ 8·4	$\frac{1}{2}$	-17·2	$\frac{1}{2}$	.....		.....		.....		.....	
Weighted mean	- 11·80		+ 8·52		- 9·37		+ 12·27		+ 47·16		.....		.....	
$V_s$	+ 28·66		+ 15·20		+ 23·24		+ 22·66		- 27·91		.....		.....	
$V_d$	+ .18		- .02		.00		+ .11		- .20		.....		.....	
Curv.	- .28		- .28		- .28		- .28		- .28		.....		.....	
Radial Velocity	+ 17·0		+ 23·4		+ 13·6		+ 34·8		+ 18·8		.....		.....	

BOSS 678

(1900,  $\alpha = 2^h 53^m \cdot 7$ ,  $\delta = +51^\circ 57'$ , mag. 5.42, type B5)

The star's spectrum consists of broad lines of hydrogen and traces of the magnesium  $\lambda 4481$ . The probable error of measurement is high, and the values given below may be as much as 10 km. in error. The  $H_\gamma$  line, however, is fairly uniform throughout its breadth and is thus fairly reliable, and judging by it alone there would seem to be good reason for considering the velocity of the star to be variable.

Plate	Date, G.M.T.	Number of lines	Velocity
8335.....	1917, Nov. 5.685	2	+11
8347.....	" 12.568	2	-44
8668.....	1918, Oct. 21.649	4	+ 3
8688.....	Dec. 9.638	3	$\pm 0$
8706.....	1919, Jan. 27.552	3	+ 4

MEASURES OF BOSS 678

$\lambda$	8335		8347		8668		8688		8706					
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4861.527					-37.8	$\frac{1}{4}$	-20.9	$\frac{1}{8}$	+30.2	$\frac{1}{4}$				
4340.634	+ 9.8	$\frac{1}{4}$	-43.0	$\frac{1}{4}$	+ 6.0	$\frac{3}{4}$	+13.5	$\frac{1}{2}$	+22.1	$\frac{3}{4}$				
4101.890	+ 4.6	$\frac{1}{2}$	-53.3	$\frac{1}{8}$	- 9.0	$\frac{1}{2}$	+19.7	$\frac{1}{8}$	+36.3	$\frac{1}{2}$				
3970.177					-17.6	$\frac{1}{2}$								
Weighted mean	+ 6.33		- 46.43		- 9.10		+ 8.80		+ 28.19					
$V_a$	+ 5.08		+ 3.00		+ 12.03		- 8.58		- 23.48					
$V_d$	+ .03		+ .14		+ .11		- .03		- .09					
Curv.	- .28		- .28		- .28		- .28		- .28					
Radial Velocity	+ 11.2		- 43.6		+ 2.8		- 0.1		+ 4.3					

## BOSS 744

(1900,  $\alpha=3^h 12^m \cdot 0$ ,  $\delta=+49^\circ 43'$ , mag. 5.08, type B3)

The hydrogen lines in the spectrum of this star are about 10 angstroms wide and thus there is considerable uncertainty in the measures. The lines  $\lambda 4471$  and  $\lambda 4026$  are the only ones seen of the helium series and they are very faint. It is not likely that the range shown in the measures represents a real variation in velocity.

Plate	Date, G.M.T.	Number of lines	Velocity
8331.....	1917, Nov. 4.787	4	- 6
8339.....	" 7.779	6	-32
8678.....	1918, Nov. 5.659	4	- 1
8689.....	Dec. 9.694	3	-16

## MEASURES OF BOSS 744

$\lambda$	8331		8339		8678		8689							
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4861.527	- 2.4	$\frac{1}{2}$	-46.5	$\frac{1}{4}$	+22.3	$\frac{1}{2}$	-39.3	$\frac{1}{2}$						
4471.676	-46.0	$\frac{1}{8}$	-56.6	$\frac{1}{4}$	-18.7	$\frac{1}{2}$								
4340.634	+ 1.1	$\frac{1}{4}$	-25.3	$\frac{1}{2}$	-23.8	$\frac{1}{2}$	- 5.5	$\frac{1}{2}$						
4101.890	-21.5	$\frac{1}{2}$	-14.7	$\frac{1}{8}$	+ 3.8	$\frac{1}{2}$	+43.8	$\frac{1}{8}$						
4026.352			-32.7	$\frac{1}{8}$										
3970.177			-64.8	$\frac{1}{8}$										
Weighted mean	- 13.09		- 38.11		- 8.00		- 8.12							
$V_a$	+ 7.55		+ 6.25		+ 7.28		- 7.88							
$V_d$	- .09		- .09		+ .09		- .09							
Curv.	- .28		- .28		- .28		- .28							
Radial Velocity	- 5.9		- 32.2		- 0.9		- 16.4							

BOSS 783

(1900,  $\alpha = 3^h 21^m \cdot 7$ ,  $\delta = +49^\circ 30'$ , mag. 5.64, type B5)

The hydrogen lines are from 10 to 15 angstroms broad and consequently the velocities obtained are only approximate. If the star is not a binary, the velocities should be about zero as it is one of the stars in the Taurus cluster. No results have hitherto been published.

Plate	Date, G.M.T.	Velocity	Lines
8357.....	1917, Nov. 14.661	+25	1
8694.....	1918, Dec. 26.564	+29	3

BOSS 839 or 40 PERSEI

(1900,  $\alpha = 3^h 36^m \cdot 0$ ,  $\delta = +33^\circ 39'$ , mag 5.04, type B2)

The measures on the 7 plates secured here are sufficiently reliable to show that this star has a variable radial velocity. A line of unknown origin, whose wave-length has been assumed as 4070.118, is quite sharp on some of the plates.

Plate	Date, G.M.T.	Velocity	Lines
8338.....	1917, Nov. 7.663	+23.2	9
8428.....	1918, Jan. 23.490	+23.7	6
8693.....	Dec. 18.719	+ 3.6	7
8696.....	" 26.668	- 8.8	6
8698.....	1919, Jan. 6.506	+33.7	4
8705.....	" 27.484	- 8.8	7
8711.....	Feb. 10.631	- 1.7	6

## MEASURES OF BOSS 839

$\lambda$	8338		8428		8693		8696		8698		8705		8711	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4861.527	- 7.2	$\frac{1}{2}$	.....	.....	+27.8	$\frac{1}{2}$	+22.3	$\frac{1}{2}$	.....	.....	+17.8	$\frac{1}{2}$	+20.2	$\frac{1}{2}$
4471.676	+19.6	$\frac{1}{2}$	+48.3	$\frac{1}{2}$	+12.4	$\frac{1}{2}$	+13.0	$\frac{1}{2}$	+63.4	$\frac{1}{2}$	+33.8	$\frac{1}{2}$	24.8	$\frac{1}{2}$
4388.100	+42.1	$\frac{1}{2}$	63.1	$\frac{1}{2}$	+ 9.6	$\frac{1}{2}$	+ 1.4	$\frac{1}{2}$	77.1	$\frac{1}{2}$	+27.1	$\frac{1}{2}$	13.1	$\frac{1}{2}$
4340.634	+17.9	$\frac{1}{2}$	66.9	$\frac{1}{2}$	+16.4	$\frac{1}{2}$	+13.1	$\frac{1}{2}$	33.0	$\frac{1}{2}$	+17.6	$\frac{1}{2}$	41.4	$\frac{1}{2}$
4143.928	+ 2.2	$\frac{1}{2}$	47.1	$\frac{1}{2}$	+27.4	$\frac{1}{2}$	- 7.1	$\frac{1}{2}$	.....	.....	.....	.....	24.7	$\frac{1}{2}$
4101.890	+38.6	$\frac{1}{2}$	.....	.....	+23.4	$\frac{1}{2}$	.....	.....	+37.2	$\frac{1}{2}$	+10.7	$\frac{1}{2}$	+30.8	$\frac{1}{2}$
4070.118	+23.1	1	37.0	$\frac{1}{2}$	.....	.....	+ 6.8	$\frac{1}{2}$	.....	.....	- 1.3	$\frac{1}{2}$	.....	.....
4026.352	- 3.4	$\frac{1}{2}$	+52.8	$\frac{1}{2}$	- 0.8	$\frac{1}{2}$	.....	.....	.....	.....	+23.6	$\frac{1}{2}$	.....	.....
3933.825	- 2.9	$\frac{1}{2}$	.....	.....	.....	.....	.....	.....	.....	.....	.....	.....	.....	.....
Weighted mean	+ 16.57		+ 50.04		+ 17.24		+ 8.36		+ 54.89		+ 18.40		+ 27.65	
$V_s$	+ 7.05		- 26.11		- 13.23		- 16.74		- 20.99		- 26.92		- 28.89	
$V_d$	- .11		+ .05		- .14		- .12		+ .10		+ .03		- .20	
Curv.	- .28		- .28		- .28		- .28		- .28		- .28		- .28	
Radial Velocity	+ 23.2		+ 23.7		+ 3.6		- 8.8		+ 33.7		- 8.8		- 1.7	

42 PERSEI

(1900,  $\alpha = 3^h 43^m \cdot 2$ ,  $\delta = +32^\circ 48'$ , mag. 5.10, type A)

This star was announced a binary by Adams in *Astrophysical Journal*, XXXV, 174, from 4 plates in 1911. Our measures indicate that the period is short.

Plate	Date, G.M.T.	Velocity	Lines
6311.....	1914, Aug. 25.870	-44	7
6459.....	Oct. 1.802	-28	6
6468.....	" 2.744	+19	6
6582.....	Nov. 27.697	-50	4

MEASURES OF 42 PERSEI

$\lambda$	6311		6459		6468		6582							
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4549	-59.8	$\frac{1}{2}$	-49.6	$\frac{1}{2}$	-9.5	$\frac{2}{3}$								
4534	68.0	$\frac{1}{2}$												
4501			46.0	$\frac{1}{2}$										
4481	63.4	$\frac{1}{2}$	54.4	$\frac{1}{2}$	-5.6	$\frac{1}{2}$	-40.3	$\frac{1}{2}$						
4340	73.1	$\frac{1}{2}$	43.3	$\frac{2}{3}$	-0.8	$\frac{1}{2}$	47.5	1						
4233			47.2	$\frac{1}{2}$										
4202							50.0	1						
4143					-23.6	$\frac{1}{2}$								
4101	68.3	$\frac{1}{2}$												
4063					+16.2	$\frac{1}{2}$								
4045	76.4	$\frac{1}{2}$												
3933	-89.0	$\frac{1}{2}$	-66.6	$\frac{1}{2}$	-2.9	$\frac{1}{2}$	-48.7	$\frac{2}{3}$						
Weighted mean	-72.73		-51.11		-3.24		-47.42							
$V_a$	+28.71		+22.86		+22.57		-2.29							
$V_d$	+ .11		+ .07		+ .14		- .02							
Curv.	- .28		- .28		- .28		- .28							
Radial Velocity	-44.2		-28.5		+19.2		-50.0							

## BOSS 898

(1900,  $\alpha = 3^{\text{h}} 48^{\text{m}} \cdot 7$ ,  $\delta = +47^{\circ} 35'$ , mag. 5.34, type B5)

The helium  $\lambda 4471$  and  $H_{\gamma}$  are the only measurable lines on our plates, both being very broad and ill-defined. The star belongs to the Taurus moving cluster and should have a velocity around zero.

Plate	Date	Velocity	Lines
8348.....	1917, Nov. 12.626	-16	2
8363.....	" 26.657	-14	2

## BOSS 947 or 48 PERSEI

(1900,  $\alpha = 4^{\text{h}} 01^{\text{m}} \cdot 4$ ,  $\delta = +47^{\circ} 27'$ , mag. 4.03, type B3)

This spectrum was described by Frost in *Astrophysical Journal*, XVIII, 389, 1903, as having bright hydrogen lines on broader absorption bands.  $H_{\beta}$  and  $H_{\gamma}$  were doubly bright, while  $H_{\delta}$  only faintly visible. Adams and Lasby in *Publications of the Astronomical Society of the Pacific*, 23, 240, record  $H_{\beta}$  and  $H_{\gamma}$  as being bright, presumably single as there is no mention of components. Merrill in *Lick Observatory Bulletin*, 237, describes 5 plates made in 1912, in which the emission showed sometimes as double and sometimes as single on the absorption bands. He gives a velocity,  $+7.1$  km. per sec., for  $H_{\beta}$  emission on the plate of August 21st.

On the 5 plates secured here the emission never occurs in the double form.  $H_{\beta}$  is always present, while  $H_{\gamma}$  is absent on the third and fourth plates and very dim on the fifth plate. From measures made upon the emission lines, there seems no doubt of a real variation in velocity of the hydrogen envelope giving rise to them. Whether this can be explained through orbital motion or not cannot be stated, but the star is worthy of further investigation. The probable error for the absorption bands is so high that the velocities given, while approximate, are not to be considered as showing a variation. The absorption bands on plate 8716 are unusually faint.

Plate	Date, G.M.T.	Absorption		Emission	
		Velocity	Lines	Velocity	Lines
8336.....	1917, Nov. 5.723	+21	3	+ 5	2
8343.....	" 9.631			-11	2
8387.....	Dec. 22.510	- 8	4	+19	1
8716.....	1919, Feb. 17.607	+10	4	-30	1
8718.....	" 24.570	+ 2	3	-10	1

MEASURES OF BOSS 947

$\lambda$	8336		8343		8387		8716		8718					
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4471.676					-30.8	$\frac{1}{2}$	+44.3	$\frac{1}{2}$						
4388.100	+ 2.5	$\frac{1}{2}$												
4143.928					+34.8	$\frac{1}{2}$	40.3	$\frac{1}{2}$	+24.7	$\frac{1}{2}$				
4101.890	14.1	$\frac{1}{2}$			+26.8	$\frac{1}{2}$								
4026.352	+13.6	$\frac{1}{2}$			- 7.8	$\frac{1}{2}$	25.2	$\frac{1}{2}$	21.9	$\frac{1}{2}$				
3933.825							+36.6	1	34.6	$\frac{1}{2}$				
4861.527 em	- 5.0	1	-16.7	$\frac{1}{2}$	+29.5	1	- 3.2	$\frac{1}{2}$	+17.0	$\frac{1}{2}$				
4340.634 em	- 7.2	$\frac{1}{2}$	-26.7	$\frac{1}{2}$										
Weighted mean	+ 10.20				+ 2.86		+ 36.60		+ 28.95					
$V_a$	+ 11.08		+ 9.40		- 10.36		- 26.55		- 26.92					
$V_d$	+ .04		+ .13		+ .13		- .18		- .13					
Curv.	- .28		- .28		- .28		- .28		- .28					
Radial Velocity	+ 21.2				- 7.6		+ 9.6		+ 1.6					
Emission	+ 5.1		- 10.8		+ 19.0		- 30.2		- 10.3					

$\psi$  ORIONIS

This star ( $\alpha = 5^h 22^m, \delta = +3^\circ 01'$ ) was announced a binary by Frost and Adams in 1903, and its orbit published by Plaskett in *Astrophysical Journal*, XXVIII, 266, 1908. Some of the plates of this star and a number of those of  $\iota$ Orionis were measured by the writer, who noted cases where the *K*-line of calcium was discrepant and where lines appeared to be double. For the purposes of the papers on "*H* and *K* lines of Calcium" and "Secondary Disturbances" given before the Ottawa Section of the R.A.S.C. sometime about 1910, the writer remeasured all the plates of these two stars where the *K*-line could be seen and where the lines were double. As these results were never published it has been thought well to give them here. In the table immediately following, the velocity of the system, + 12.0 km. per sec., has been subtracted from the measured velocities so as to get the relative orbital motion from which is deduced the ratios of the masses given in the fifth column. The second column simply gives Plaskett's velocities with 12 km. subtracted from each. From the last two columns the mean ratio of the masses is 0.63.

## VELOCITIES OF COMPONENT STARS

Plate	Velocity		Lines	m/m
	Primary	Secondary		
1158.....	-157.5	+222.9	4	.71
1183.....	+127.1	-186.7	6	.68
1208.....	+133.9	-242.3	4	.55
1209.....	+136.0	-229.0	1	.59
1238.....	-126.5	+224.4	5	.56
1271.....	+127.9	-240.7	2	.53
1304.....	+139.6	-226.8	1	.57
1312.....	+133.3	-202.5	4	.66
1317.....	+144.0	-198.3	4	.72
1319.....	+128.5	-180.8	5	.71
1336.....	+ 91.3	-168.7	3	.54
1347.....	-147.1	+245.5	4	.60
1395.....	-147.0	+279.8	3	.53

The *K*-line seems to vary, though the range is less than that for the other lines. A check plate, No. 4852 taken Feb. 23.502, 1912, agrees in the case of the main lines with the curve within the limits of the probable error, and its value for *K* is added in the table following. To clear the fractions, the weights assigned at the time of measurement are multiplied by 4 in the table following, in which the velocities quoted are simply those as measured relative to the sun. The error of measurement is large, but it should be within 15 or 20 km.

## K-LINE VELOCITY

Plate	Velocity	Weight	Plate	Velocity	Weight
1138.....	+82	2	1239.....	- 9	2
1183.....	+14	1	1296.....	- 8	2
1209.....	-25	2	1312.....	+ 4	2
1214.....	+55	1	1317.....	+53	2
1215.....	+55	2	1333.....	+33	4
1220.....	+23	2	1347.....	-94	2
1221.....	+81	2	1349.....	-65	1
1233.....	+16	1	4852.....	+12	4

ORIONIS

The orbit of this star ( $\alpha = 5^h 30^m$ ,  $\delta = -5^\circ 59'$ ) was published by Plaskett in *Astrophysical Journal*, volume XXX, 373, 1909. On a fine-grained plate, No. 4847 taken Feb. 20·469, 1912, five lines belonging to the secondary component were measured and from these a ratio of the masses of 0·58 was obtained. Unlike  $\psi$  Orionis the K-line velocity seems constant, or at least of very small range. The weighted mean velocity for 30 plates, using the wave-length 3933·825, is + 30·1 km. per second, agreeing closely with that for the velocity of the system, + 21·3 km. per second. Here, as in  $\psi$  Orionis, the weights published are four times those given at measurement.

VELOCITY K-LINE

Plate	Velocity	Weight	Plate	Velocity	Weight
1076.....	+12·6	2	1162.....	+16·1	2
1077.....	39·6	3	1170.....	23·2	4
1078.....	43·6	3	1190.....	24·4	4
1112.....	36·8	3	1194.....	35·6	2
1116.....	12·0	1	1201.....	55·2	1
1119.....	28·4	1	1207.....	17·2	1
1120.....	21·2	2	1212.....	43·7	4
1122.....	34·6	1	1213.....	43·6	3
1123.....	36·0	4	1219.....	37·7	2
1124.....	41·3	2	1263.....	19·9	3
1125.....	32·9	4	1266.....	11·0	2
1126.....	9·1	1	1275.....	26·7	2
1136.....	20·6	2	1277.....	18·2	2
1143.....	26·8	2	1278.....	38·4	2
1148.....	+27·6	4	4847.....	+32·9	2

30 CANIS MAJORIS

(1900,  $\alpha = 7^h 14^m \cdot 5$ ,  $\delta = -24^\circ 47'$ , mag. 4·40, type 0e5)

This star was announced a spectroscopic binary by Frost in a footnote in *Astrophysical Journal*, XXIII, page 265. Later Lee gave the velocities of five plates in *Astrophysical Journal* XXXIX, page 45. Four measures are also given by Campbell in *Lick Observatory Bulletin*, 199. The star would be a profitable and easy one to work up here but for its southern declination.

Plate	Date, G.M.T.	Velocity	Lines
8383.....	1917, Dec. 11·833	+32	2
8408.....	1918, Jan. 3·733	+71	2
8453.....	Feb. 20·604	+ 1	3

## MEASURES OF 30 CANIS MAJORIS

$\lambda$	8383		8408		8453									
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4861.527					+28.6	$\frac{1}{2}$								
4471.676	+25.7	$\frac{1}{2}$	+76.2	$\frac{1}{2}$	15.4	$\frac{1}{2}$								
4340.634	+ 8.6	$\frac{1}{2}$	+49.7	$\frac{1}{2}$	+ 6.8	$\frac{1}{2}$								
Weighted mean	+ 20.00		+ 67.37		+ 13.07									
$V_a$	+ 11.96		+ 4.33		- 12.22									
$V_d$	- .09		.00		- .02									
Curv.	- .28		- .28		- .28									
Radial Velocity	+ 31.6		+ 71.4		+ 0.6									

## BOSS 2381

(1900,  $\alpha = 8^h 46^m \cdot 6$ ,  $\delta = - 6^\circ 49'$ , mag. 5.60, type A2)

This star was announced as a spectroscopic binary by Adams in *Publications of the Astronomical Society of the Pacific*, XXVI, 261.

Our two measures follow:—

Plate	Date, G.M.T.	Velocity	Lines
7972.....	1917, Dec. 29.807	+49	9
8000.....	1918, Jan. 16.773	+33	4

MEASURES OF BOSS 2381

$\lambda$	7972		8000											
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4584.018			+ 1.6	$\frac{1}{2}$										
4572.156	+42.3	$\frac{1}{2}$												
4563.939	25.5	$\frac{1}{2}$												
4549.766	30.0	$\frac{3}{4}$	24.6	$\frac{1}{2}$										
4481.400	43.9	1	35.5	$\frac{3}{4}$										
4340.634	16.8	$\frac{1}{2}$												
4325.939	12.2	$\frac{1}{2}$												
4271.760	40.4	$\frac{1}{2}$	+ 8.2	$\frac{1}{2}$										
4233.328	40.8	$\frac{1}{2}$												
4045.975	+28.0	$\frac{1}{2}$												
Weighted mean	+ 32.04		+ 23.66											
$V_a$	+ 17.06		+ 9.48											
$V_d$	.00		- .04											
Curv.	- .28		- .28											
Radial Velocity	+ 48.8		+ 32.8											

36 LYNCIS

(1900,  $\alpha = 9^h 07^m \cdot 3$ ,  $\delta = + 43^\circ 38'$ , mag. 5.30, type B8)

While no range is shown in our measures, Jordan at the Allegheny Observatory gets a range of 20 km. on 7 plates—from + 4 to + 24—so the star is probably a spectroscopic binary as the lines are narrow and well adapted for measurement.

Plate	Date, G.M.T.	Velocity	Lines	Weight
6939	1915, April 20.603	+13.8	6	4
6953	" 26.612	+19.5	3	2
6978	May 10.577	+17.5	9	5
6993	" 14.575	+15.2	6	4
8511	1918, April 24.592	+10.7	4	2

## MEASURES OF 36 LYNCIS

$\lambda$	6939		6953		6978		6993		8511					
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4572·156					+63·1	$\frac{1}{2}$								
4549·766	+31·6	1	+47·9	$\frac{1}{2}$	41·4	$\frac{2}{4}$	+34·4	$\frac{1}{2}$	+45·6	$\frac{1}{2}$				
4481·400	38·5	$\frac{2}{4}$	41·2	$\frac{1}{2}$	30·0	$\frac{2}{4}$	36·4	$\frac{1}{2}$	26·4	$\frac{1}{2}$				
4340·634	50·6	$\frac{2}{4}$			38·0	1	45·0	1	43·6	$\frac{1}{2}$				
4338·084							41·5	$\frac{1}{2}$						
4233·328					55·3	$\frac{1}{2}$			+32·7	$\frac{1}{2}$				
4131·047					52·5	$\frac{1}{2}$								
4128·211	42·0	$\frac{1}{2}$	+50·4	$\frac{1}{2}$	54·3	$\frac{1}{2}$								
4101·890	44·7	$\frac{1}{2}$			46·6	$\frac{1}{2}$	46·4	$\frac{2}{4}$						
3933·825	+42·2	$\frac{1}{2}$			+36·8	$\frac{1}{2}$	+32·7	$\frac{1}{2}$						
Weighted mean	+ 40·70		+ 46·50		+ 43·80		+ 41·03		+ 37·73					
$V_a$	- 26·44		- 26·59		- 25·89		- 25·42		- 26·58					
$V_d$	- .13		- .17		- .17		- .18		- .13					
Curv.	- .28		- .28		- .28		- .28		- .28					
Radial Velocity	+ 13·8		+ 19·5		+ 17·5		+ 15·2		+ 10·7					

ζ LEONIS

(1900,  $\alpha = 10^h 11^m \cdot 1$ ,  $\delta = + 23^\circ 55'$ , mag. 3.65, type F)

This star was announced a binary by Campbell in *Lick Observatory Bulletin*, 199.  
Our measures follow:

Plate	Date, G.M.T.	Velocity	Lines
5325.....	1913, Jan. 27.734	-12.7	10
5331.....	“ 28.722	-16.3	9
5342.....	Feb. 3.857	- 8.5	5

MEASURES OF ζ LEONIS

λ	5325		5331		5342									
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4395.286	-26.0	$\frac{1}{2}$												
4352.006	14.7	$\frac{1}{2}$	-36.4	$\frac{1}{2}$										
4340.634	14.4	$\frac{3}{2}$	11.6	$\frac{1}{2}$	- 3.7	$\frac{1}{2}$								
4325.939			24.8	$\frac{1}{2}$										
4271.760	18.9	$\frac{1}{2}$	29.3	$\frac{1}{2}$	-11.3	$\frac{1}{2}$								
4198.494	4.8	$\frac{1}{2}$	0.8	$\frac{1}{2}$										
4143.928	40.5	$\frac{1}{2}$	36.2	$\frac{1}{2}$	-30.4	$\frac{1}{2}$								
4101.890	20.7	$\frac{2}{2}$	25.0	$\frac{1}{2}$										
4063.756	39.1	$\frac{1}{2}$			-29.0	$\frac{1}{2}$								
4045.975	24.5	$\frac{1}{2}$	36.2	$\frac{2}{2}$										
4005.597	-18.6	$\frac{1}{2}$	-21.4	$\frac{1}{2}$	+ 3.7	$\frac{1}{2}$								
Weighted mean	- 22.11		- 25.24		- 14.12									
V <sub>a</sub>	+ 9.64		+ 9.13		+ 6.09									
V <sub>d</sub>	+ .07		+ .08		- .17									
Curv.	- .28		- .28		- .28									
Radial Velocity	- 12.7		- 16.3		- 8.5									

$\pi$  8 VIRGINIS(1900,  $\alpha = 11^{\text{h}} 55^{\text{m}}.7$ ,  $\delta = + 7^{\circ} 10'$ , mag. 4.57, type A3)

This star was announced a spectroscopic binary by Albrecht in *Lick Observatory Bulletin*, V, 175, from three plates in 1909 giving a range from + 18 to - 21. In *Astrophysical Journal*, vol. XXXIX, 46, Lee gives the measures of four plates made in 1906, 1907 and 1909 showing a variation from + 6 to - 33. Two plates were made here in 1910 and their measures published in the Chief Astronomer's report for 1911. Their velocities are slightly changed in the table below to agree with the different wave-lengths used in the recent measures. The thirteen additional plates secured all give quite negative velocities, and it would appear that the time when positive velocities are possible is very short relative to the whole period, a condition which would obtain if the eccentricity were fairly high and  $\omega$  around zero. The velocity of the system would seem to be about 20 km. per sec. negative. The lines are only fair for measurement on our single-prism plates. The measures follow.

Plate	Date, G.M.T.	Velocity	Lines
3349.....	1910, Mar. 18.868	-26.0	3
3383.....	April 11.792	21.8	3
8066.....	1917, Feb. 18.806	16.9	12
8078.....	" 27.830	26.4	10
8085.....	Mar. 1.751	35.8	13
8094.....	" 2.810	29.8	9
8133.....	" 30.727	26.4	6
8136.....	April 3.601	35.2	6
8139.....	" 8.590	35.7	9
8143.....	" 10.642	44.2	8
8147.....	" 16.753	30.7	9
8152.....	" 22.578	30.2	10
8155.....	" 23.702	42.5	7
8159.....	" 24.676	30.7	7
8166.....	May 13.725	-33.7	4

MEASURES OF  $\pi$  8 VIRGINIS

$\lambda$	3349		3383		8066		8078		8085		8094		8133	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4584.018	-31.9	$\frac{1}{2}$												
4549.743	18.2	$\frac{1}{2}$	-29.6	$\frac{1}{2}$	-39.8	$\frac{1}{2}$	-35.4	$\frac{1}{2}$	-46.6	$\frac{1}{2}$	-42.6	$\frac{1}{2}$	-45.3	$\frac{1}{2}$
4534.158									65.0	$\frac{1}{2}$				
4515.508											25.8	$\frac{1}{2}$		
4501.417									29.6	$\frac{1}{2}$	48.7	$\frac{1}{2}$		
4481.477	-27.5	$\frac{1}{2}$			26.5	$\frac{1}{2}$	40.8	$\frac{1}{2}$	44.7	$\frac{1}{2}$	43.4	$\frac{1}{2}$		
4395.155									55.8	$\frac{1}{2}$	38.2	$\frac{1}{2}$		
4340.645			0.0	$\frac{1}{2}$	47.1	$\frac{1}{2}$	46.3	$\frac{1}{2}$	53.7	$\frac{1}{2}$	41.4	$\frac{1}{2}$	-37.7	$\frac{1}{2}$
4325.698					32.2	$\frac{1}{2}$	22.9	$\frac{1}{2}$					+ 3.7	$\frac{1}{2}$
4307.974									41.0	$\frac{1}{2}$			-22.9	$\frac{1}{2}$
4290.053					36.5	$\frac{1}{2}$	10.0	$\frac{1}{2}$			17.5	$\frac{1}{2}$		
4282.584													+14.3	$\frac{1}{2}$
4271.765					17.2	$\frac{1}{2}$			22.6	$\frac{1}{2}$				
4236.000							29.8	$\frac{1}{2}$						
4233.425					39.2	$\frac{1}{2}$	62.4	$\frac{1}{2}$	48.4	$\frac{1}{2}$	34.5	$\frac{1}{2}$	-34.3	$\frac{1}{2}$
4227.107							22.3	$\frac{1}{2}$						
4215.733											-35.5	$\frac{1}{2}$		
4198.677					8.4	$\frac{1}{2}$			19.2	$\frac{1}{2}$				
4143.839			- 3.1	$\frac{1}{2}$	29.1	$\frac{1}{2}$	36.8	$\frac{1}{2}$	65.3	$\frac{1}{2}$				
4071.865									37.8	$\frac{1}{2}$				
4063.730					40.6	$\frac{1}{2}$								
4045.940					23.4	$\frac{1}{2}$			-41.2	$\frac{1}{2}$				
4005.414					-23.2	$\frac{1}{2}$	-45.6	$\frac{1}{2}$						
Weighted mean	- 25.13		- 9.19		- 30.27		- 35.23		- 43.74		- 37.23		- 20.55	
$V_a$	- 0.35		- 12.10		+ 13.60		+ 9.24		+ 8.20		+ 7.76		- 5.44	
$V_s$	- .26		- .21		- .02		- .12		+ .02		- .09		- .09	
Curv.	- .28		- .28		- .28		- .28		- .28		- .28		- .28	
Radial Velocity	- 26.0		- 21.8		- 16.9		- 26.4		- 35.8		- 29.8		- 26.4	

MEASURES OF  $\pi$  8 VIRGINIS—*Concluded*

$\lambda$	8136		8139		8143		8147		8152		8155		8159	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4584.018							-11.5	$\frac{1}{2}$	-22.7	$\frac{1}{2}$				
4572.190	-21.1	$\frac{1}{2}$	-11.1	$\frac{1}{2}$									-9.6	$\frac{1}{2}$
4563.939					-49.7	$\frac{1}{2}$								
4549.743	19.1	$\frac{1}{2}$			34.6	$\frac{1}{2}$	24.5	$\frac{1}{2}$	-8.7	$\frac{1}{2}$	-27.5	$\frac{1}{2}$	+17.6	$\frac{1}{2}$
4481.477	23.5	$\frac{1}{2}$	26.4	$\frac{1}{2}$	32.7	$\frac{1}{2}$	23.9	$\frac{1}{2}$	-10.2	$\frac{1}{2}$	33.9	$\frac{1}{2}$	-30.2	$\frac{1}{2}$
4351.977							14.6	$\frac{1}{2}$						
4340.645	5.2	$\frac{1}{2}$	13.2	$\frac{1}{2}$	33.8	$\frac{1}{2}$	14.4	$\frac{1}{2}$	-15.0	$\frac{1}{2}$	10.6	$\frac{1}{2}$	-24.5	$\frac{1}{2}$
4325.698					21.2	$\frac{1}{2}$	31.6	$\frac{1}{2}$			18.3	$\frac{1}{2}$	-27.0	$\frac{1}{2}$
4307.974					38.6	$\frac{1}{2}$			-7.8	$\frac{1}{2}$	29.2	$\frac{1}{2}$	+12.4	$\frac{1}{2}$
4300.211	52.1	$\frac{1}{2}$												
4294.359											22.9	$\frac{1}{2}$		
4290.053			34.0	$\frac{1}{2}$										
4282.584									+10.6	$\frac{1}{2}$				
4246.996									-5.2	$\frac{1}{2}$				
4236.000											-30.8	$\frac{1}{2}$		
4233.425			18.2	$\frac{1}{2}$			10.6	$\frac{1}{2}$	-12.2	$\frac{1}{2}$			-16.1	$\frac{1}{2}$
4227.107			38.4	$\frac{1}{2}$										
4215.733			31.4	$\frac{1}{2}$										
4101.898			21.4	$\frac{1}{2}$	23.9	$\frac{1}{2}$								
4071.865					-21.7	$\frac{1}{2}$	0.0	$\frac{1}{2}$	-37.9	$\frac{1}{2}$				
4045.940	-26.9	$\frac{1}{2}$	-38.5	$\frac{1}{2}$			-18.1	$\frac{1}{2}$	-23.2	$\frac{1}{2}$				
Weighted mean	-26.67		-24.84		-32.20		-15.82		-13.10		-24.74		-12.57	
$V_s$	-8.36		-10.73		-11.69		-14.43		-16.90		-17.35		-17.74	
$V_d$	+ .14		+ .14		+ .03		- .21		+ .09		- .16		- .11	
Curv.	- .28		- .28		- .28		- .28		- .28		- .28		- .28	
Radial Velocity	-35.2		-35.7		-44.2		-30.7		-30.2		-42.5		-30.7	

16 COMÆ

(1900,  $\alpha = 12^h 22^m \cdot 0$ ,  $\delta = + 27^\circ 22'$ , mag. 5.04, type A2)

The hydrogen lines and the calcium *K* are very strong in this spectrum, the magnesium  $\lambda 4481$  fairly strong, while numerous other metallic lines are faint and hard to measure. The results, which are the means of two independent measures, do not establish a variable velocity.

Plate	Date, G.M.T.	Velocity	Lines
6941.....	1915, April 20.666	$\pm 0$	11
6974.....	May 9.704	- 4	11
6995.....	" 14.676	-13	14
7062.....	June 17.594	- 4	15

MEASURES OF 16 COMÆ

$\lambda$	6941		6974		6995		7062							
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4584.018							+15.7	$\frac{1}{2}$						
4572.190			+10.4	$\frac{1}{2}$										
4549.743	+ 2.1	$\frac{1}{2}$	+17.6	$\frac{1}{2}$	+ 2.0	$\frac{1}{2}$	+26.2	$\frac{3}{4}$						
4534.158							+34.1	$\frac{1}{2}$						
4508.445							- 6.9	$\frac{1}{2}$						
4481.477	+14.0	$\frac{1}{2}$	+19.2	$\frac{1}{2}$	+17.2	$\frac{1}{2}$	+31.8	$\frac{3}{4}$						
4415.345							+39.6	$\frac{1}{2}$						
4404.861							+30.8	$\frac{1}{2}$						
4351.977	+43.0	$\frac{1}{4}$												
4340.645	+26.0	1	+ 3.1	$\frac{1}{2}$	- 0.8	$\frac{1}{2}$	+34.3	$\frac{1}{2}$						
4325.698					+ 4.8	$\frac{1}{2}$	+30.3	$\frac{3}{4}$						
4307.974	+ 5.8	$\frac{1}{2}$												
4250.586							+33.3	$\frac{1}{2}$						
4233.425	+32.5	$\frac{1}{2}$	+21.0	$\frac{1}{2}$	+25.6	$\frac{1}{2}$								
4227.107					-21.6	$\frac{1}{2}$								
4202.192			+28.1	$\frac{3}{4}$	+31.9	$\frac{1}{4}$								
4198.677			+28.4	$\frac{1}{2}$	+13.8	$\frac{1}{2}$								
4143.839							- 4.5	$\frac{3}{4}$						
4128.211	+36.5	$\frac{1}{4}$			+18.6	$\frac{1}{2}$								
4101.898	- 1.6	$\frac{1}{2}$	+ 5.6	$\frac{1}{2}$	+ 2.4	$\frac{1}{2}$	+40.0	$\frac{1}{4}$						
4071.865					+17.2	$\frac{1}{2}$	+25.9	$\frac{1}{2}$						
4063.730	- 2.8	$\frac{1}{2}$	+18.1	$\frac{1}{2}$	+ 4.5	$\frac{1}{2}$	- 0.9	$\frac{1}{2}$						
4045.940	+26.1	$\frac{1}{2}$	+29.0	$\frac{1}{2}$	+ 9.0	$\frac{1}{2}$								
4005.414							+20.1	$\frac{1}{2}$						
3933.825	+10.6	$\frac{1}{2}$	- 3.8	$\frac{1}{4}$	- 4.0	$\frac{1}{4}$								
Weighted mean	+ 15.73		+ 17.89		+ 10.24		+ 22.96							
$V_a$	- 15.36		- 21.28		- 22.48		- 26.04							
$V_d$	- .11		- .16		- .16		- .16							
Curv.	- .28		- .28		- .28		- .28							
Radial Velocity	$\pm 0.0$		- 3.8		- 12.7		- 3.5							

## 12 CANUM VENATICORUM

(1900,  $\alpha = 12^{\text{h}} 51^{\text{m}} \cdot 4$ ,  $\delta = + 38^{\circ} 52'$ , mag. 5.39, type Ap)

This is the fainter of the pair No. 6313 in Burnham's General Catalogue, whose distance apart of  $20''$  seems to be constant. The stars have a common proper motion of  $0'' \cdot 257$ , and thus, if the components of the pair should differ appreciably in their radial velocities, it would suggest a binary character for at least one of them: For the brighter star Campbell quotes a velocity of  $-1.6$  km. per sec. in *Lick Observatory Bulletin*, 211, and Hnatek in *A.N.*, 197, 185, gives five velocities, the mean of which is  $+2.0$  km. per sec. Thus, a velocity in the neighbourhood of  $\pm 0$  km. per sec. may be accepted for this star. For the fainter star, here dealt with, Ludendorff gives a velocity of  $-0.3$  km. per sec., being the mean of 12 accordant plates and agreeing well with the result for the main star. Our results, however, reveal a variable velocity, as either the first or second plate, both of which have quite sharp and well measurable lines, gives a velocity too divergent from the other measures to consider it as constant.

Plate	Date, G.M.T.	Velocity	Lines
7028.....	1915, May 30.613	-14.8	16
7104.....	July 15.616	-21.0	14
8079.....	1917, Feb. 27.884	- 3.6	10
8086.....	Mar. 1.812	- 8.0	12

MEASURES OF 12 CANUM VENATICORUM (fainter)

$\lambda$	7028		7104		8079		8086							
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4584.018	- 8.9	1					-18.1	$\frac{1}{2}$						
4572.156	+29.3	$\frac{1}{2}$	- 2.0	$\frac{1}{2}$	+ 5.8	$\frac{1}{2}$	-11.3	$\frac{1}{2}$						
4563.939	+29.2	1			-14.4	$\frac{1}{2}$								
4558.827	+ 8.0	$\frac{1}{2}$												
4549.766	+10.7	$\frac{1}{2}$	- 6.5	$\frac{1}{2}$	-10.6	$\frac{1}{2}$	-20.5	$\frac{1}{2}$						
4534.139					-11.5	$\frac{1}{2}$								
4481.400					+ 4.0	$\frac{1}{2}$								
4468.663	+21.2	1	+15.0	$\frac{1}{2}$										
4415.301	- 8.0	$\frac{1}{2}$	+ 7.8	$\frac{1}{2}$										
4404.927	+ 9.7	1			-11.8	$\frac{1}{2}$	-15.9	$\frac{1}{2}$						
4395.286	+ 4.0	$\frac{1}{2}$												
4383.720					-16.6	$\frac{1}{2}$								
4352.006			- 9.6	$\frac{1}{2}$			-26.4	$\frac{1}{2}$						
4340.634			-10.3	$\frac{1}{2}$			+ 9.7	$\frac{1}{2}$						
4325.939			- 7.5	1	-19.6	$\frac{1}{2}$	- 8.0	$\frac{1}{2}$						
4308.081	+ 7.5	1	- 3.6	$\frac{1}{2}$			-16.5	$\frac{1}{2}$						
4294.301			-10.8	$\frac{1}{2}$										
4289.915							-13.2	$\frac{1}{2}$						
4271.760	+10.0	1	+ 7.3	$\frac{1}{2}$			-21.8	$\frac{1}{2}$						
4260.640	-10.6	1												
4250.616	+ 1.3	$\frac{1}{2}$	+18.6	$\frac{1}{2}$										
4236.107			- 8.8	$\frac{1}{2}$			-16.5	$\frac{1}{2}$						
4233.328	+ 6.2	$\frac{1}{2}$	+ 1.2	$\frac{1}{2}$			+13.5	$\frac{1}{2}$						
4215.668					-11.9	$\frac{1}{2}$								
4202.198	+ 1.1	1												
4143.928			- 9.4	$\frac{1}{2}$	- 4.5	$\frac{1}{2}$								
4045.975	+ 9.5	$\frac{1}{2}$												
Weighted mean	+ 7.84		- 0.95		- 9.11		- 12.84							
$V_a$	- 21.65		- 19.52		+ 5.92		+ 5.18							
$V_d$	- 0.09		- 0.22		- 0.11		- 0.02							
Curv.	- 0.28		- 0.28		- 0.28		- 0.28							
Radial Velocity	- 14.8		- 21.0		- 3.6		- 8.0							

$\tau$  93 VIRGINIS(1900,  $\alpha = 13^{\text{h}} 56^{\text{m}} \cdot 6$ ,  $\delta = + 2^{\circ} 02'$ , mag. 4.34, type A2)

The spectrum of this star consists of broad hydrogen lines and calcium *K*, while hazy ill-defined  $\lambda$  4481 and  $\lambda$  4549 are also seen. Occasionally traces of other metallic lines appear, but the measures are confined to the ones above mentioned. While the range of 46 km. obtained from the measures might suggest a real variation in the radial velocity, the writer is inclined to ascribe the greatest portion of it at least to errors of measurement. Treating it as of constant velocity, we get a value  $- 2$  km. per sec. Campbell in *L. O. B.*, 211, uses the value  $- 10$ : in his statistical treatment of the velocities of A-type stars.

Plate	Date, G.M.T.	Velocity	Lines
1342.....	1908, Feb. 21-88	- 9	5
1382.....	Mar. 4-86	+18	3
1389.....	" 9-81	+10	3
1399.....	" 16-83	- 3	3
1510.....	April 22-82	-12	3
1535.....	May 18-61	-26	3
1570.....	June 3-65	-15	3
2417.....	1909, Mar. 22-82	- 1	4
2453.....	" 31-80	- 1	2
2502.....	April 19-84	+21	2
2512.....	" 23-77	+ 2	3
2532.....	" 28-80	-20	2

MEASURES OF  $\tau$  VIRGINIS

$\lambda$	1342		1382		1389		1399		1510		1535		1570	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4861.527													-13.4	$\frac{1}{2}$
4549.766									-10.4	$\frac{1}{2}$				
4481.400	-49.8	$\frac{1}{2}$	+17.3	$\frac{1}{2}$	+18.1	$\frac{1}{2}$			-22.3	$\frac{1}{2}$				
4340.634	40.4	$\frac{1}{2}$	-12.1	$\frac{1}{2}$	- 6.8	$1\frac{1}{2}$	-28.6	$\frac{1}{2}$	+ 4.5	$\frac{1}{2}$	-16.0	$\frac{1}{2}$	+ 5.4	$\frac{1}{2}$
4101.890	28.8	$\frac{1}{2}$	+ 0.9	$\frac{1}{2}$	-31.7	$\frac{1}{2}$	11.6	$\frac{1}{2}$			13.4	$\frac{1}{2}$		
3970.177	13.6	$\frac{1}{2}$												
3933.825	-43.4	$\frac{1}{2}$					-20.2	$\frac{1}{2}$			-25.5	$\frac{1}{2}$	+27.0	$\frac{1}{2}$
Weighted mean	- 32.34		- 1.51		- 8.74		- 18.00		- 9.40		- 19.28		+ 6.10	
$V_a$	+ 23.96		+ 19.89		+ 17.97		+ 15.02		- 1.90		- 6.73		- 20.61	
$V_d$	- .04		- .04		+ .04		- .04		- .20		+ .04		- .12	
Curv.	- .28		- .28		- .28		- .28		- .28		- .28		- .28	
Radial Velocity	- 8.7		+ 18.1		+ 10.0		- 3.3		- 11.8		- 26.2		- 14.9	

MEASURES OF  $\tau$  VIRGINIS—*Concluded*

$\lambda$	2417		2453		2502		2512		2532		Vel.	Wt.	Vel.	Wt.	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.					
4861·527	- 6·6	$\frac{1}{8}$													
4549·766							+16·7	$\frac{1}{4}$							
4481·400	+ 6·5	$\frac{1}{8}$	+ 4·5	$\frac{1}{2}$	+24·3	$\frac{1}{2}$	+19·6	$\frac{1}{2}$							
4340·634	-22·6	$\frac{1}{2}$	-22·4	$\frac{1}{2}$	+21·1	$\frac{1}{2}$	- 7·3	$\frac{1}{2}$	-19·7	$\frac{1}{2}$					
4101·890	- 0·6	$\frac{1}{8}$							+11·2	$\frac{1}{8}$					
Weighted mean	- 12·89		- 9·00		+ 22·74		+ 5·42		- 13·52						
$V_s$	+ 12·45		+ 8·23		- 1·09		- 3·05		- 5·49						
$V_d$	- .04		- .08		- .22		- .10		- .23						
Curv.	- .28		- .28		- .28		- .28		- .28						
Radial Velocity	- 0·8		- 1·1		+ 21·2		+ 2·0		- 19·5						



10 SERPENTIS

(1900,  $\alpha = 15^h 23^m \cdot 6$ ,  $\delta = + 2^\circ 12'$ , mag. 5.12, type A5)

The lines in this spectrum are very diffuse and ill-defined. Plate 7728 is much underexposed, so that the star's variable velocity is not established by these measures.

	Plate	Date, G.M.T.	Velocity	n	Weight
7030.....		1915, May 30.709	-23	7	2
7082.....		July 1.551	-30	7	2
7091.....		" 9.546	-35	10	3
7728.....		1916, " 13.545	- 4	5	2

MEASURES OF 10 SERPENTIS

$\lambda$	7030		7082		7091		7728							
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4584.191							+29.7	$\frac{1}{2}$						
4549.766	-13.0	$\frac{1}{4}$	-32.6	$\frac{1}{4}$	-12.2	$\frac{1}{4}$	-11.6	$\frac{1}{4}$						
4481.400	-17.0	$\frac{1}{4}$	- 1.8	$\frac{1}{4}$	-19.2	$\frac{1}{4}$	+49.9	$\frac{1}{2}$						
4468.663	- 0.1	$\frac{1}{4}$												
4352.006			+ 1.3	$\frac{1}{4}$										
4340.634	+ 1.6	$\frac{1}{4}$	-10.0	$\frac{1}{4}$	+ 8.7	$\frac{1}{4}$								
4325.939	-19.4	$\frac{1}{4}$	-12.1	$\frac{1}{4}$	-20.6	$\frac{1}{4}$	- 1.1	$\frac{1}{2}$						
4308.081					+17.4	$\frac{1}{4}$								
4289.915					- 0.8	$\frac{1}{4}$								
4233.328	-18.4	$\frac{1}{4}$	-11.0	$\frac{1}{4}$	-17.9	$\frac{1}{4}$								
4227.010	-28.1	$\frac{1}{4}$					+21.8	$\frac{1}{4}$						
4198.494					+ 1.2	$\frac{1}{4}$								
4143.928					-22.2	$\frac{1}{4}$								
4045.975			+ 5.5	$\frac{1}{4}$	-43.3	$\frac{1}{4}$								
Weighted mean	- 13.47		- 8.66		- 10.89		+ 21.20							
$V_a$	- 9.42		- 21.34		- 23.48		- 24.55							
$V_d$	- 0.09		- 0.04		- 0.06		- 0.09							
Curv.	- 0.28		- 0.28		- 0.28		- 0.28							
Radial Velocity	- 23.3		- 30.3		- 34.7		- 3.7							



21 OPHIUCHI

(1900,  $\alpha = 16^h 46^m \cdot 4$ ,  $\delta = + 1^\circ 23'$ , mag. 5.47, type A)

The lines  $\lambda 4549$  and  $\lambda 4481$  are fairly good lines in this spectrum. Plate 7009 is underexposed.

Plate	Date, G.M.T.	Velocity	Lines	Weight
6985.....	1915, May 11.777	-27	4	2
7009.....	" 23.765	-25	3	1
7037.....	" 31.773	-33	4	1
7063.....	June 17.645	-32	6	3

MEASURES OF 21 OPHIUCHI

$\lambda$	6985		7009		7037		7063							
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4584.191			-51.2	$\frac{1}{4}$	-43.7	$\frac{1}{8}$								
4549.766	-37.1	$\frac{3}{4}$	36.0	$\frac{1}{2}$	31.8	$\frac{1}{2}$	-29.0	$\frac{1}{2}$						
4481.400	40.7	$\frac{1}{2}$	-9.4	$\frac{1}{2}$	31.7	$\frac{1}{2}$	12.9	$\frac{3}{4}$						
4340.634	33.8	$\frac{1}{4}$												
4325.939							17.9	$\frac{1}{4}$						
4308.081							23.2	$\frac{1}{4}$						
4233.328					-29.4	$\frac{1}{4}$	22.5	$\frac{1}{2}$						
4045.975	-22.8	$\frac{1}{4}$					-37.7	$\frac{3}{4}$						
Weighted mean	- 36.30		- 28.40		- 32.64		- 24.60							
$V_a$	+ 9.38		+ 4.05		+ 0.40		- 7.18							
$V_d$	.00		-.04		-.10		+ .07							
Curv.	-.28		-.28		-.28		-.28							
Radial Velocity	- 27.2		- 24.7		- 32.6		- 32.0							

## 101 HERCULIS

(1900,  $\alpha = 18^{\text{h}} 04^{\text{m}} \cdot 6$ ,  $\delta = + 20^{\circ} 01'$ , mag. 5.24, type A2)

The lines in this spectrum are excellent for measurement. No variation is indicated by the measures which follow.

Plate	Date, G.M.T.	Velocity	Lines	Weight
6976.....	1915, May 9.812	-16.6	10	10
6990.....	" 13.770	-21.8	12	12
7024.....	" 29.711	-21.3	15	12
7075.....	June 25.642	-19.7	17	15

## MEASURES OF 101 HERCULIS

$\lambda$	6976		6990		7024		7075		Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.						
4549.746	-36.2	1	-33.7	1	-30.9	1	-6.6	1						
4534.139							28.1	$\frac{1}{2}$						
4481.464	29.8	1	40.2	1	34.4	1	16.9	1						
4415.333					24.0	1	7.8	$\frac{1}{2}$						
4404.857							22.9	1						
4395.147							10.1	1						
4351.991							8.8	$\frac{1}{2}$						
4340.667	36.8	1	32.2	1										
4325.707	57.0	1	52.7	1	38.7	$\frac{1}{2}$	30.0	1						
4307.980					46.7	$\frac{1}{2}$	31.5	1						
4290.070					41.0	$\frac{1}{2}$	22.5	1						
4271.645					31.3	1	21.4	1						
4250.687							28.0	1						
4235.994					25.7	$\frac{1}{2}$	18.6	1						
4233.421	28.3	1	24.7	1	32.0	1	19.0	1						
4227.124					31.3	$\frac{1}{2}$								
4202.278					37.5	$\frac{1}{2}$								
4198.719	24.3	1	39.4	1										
4143.839	26.0	1	32.8	1	24.6	1	7.0	1						
4101.890			37.6	1										
4077.862	35.9	1	33.0	1										
4071.861							23.5	1						
4063.706			30.5	1	19.3	1								
4045.929	18.0	1	37.7	1	26.0	1	-13.8	1						
4005.402	-18.6	1	-29.2	1	-26.6	1								
Weighted mean	-31.09		-35.31		-29.94		-18.97							
$V_a$	+14.82		+13.73		+8.85		-0.60							
$V_z$	-.04		+.10		+.11		+.11							
Curv.	-.28		-.28		-.28		-.28							
Radial Velocity	-16.6		-21.8		-21.3		-19.7							



## ADDITIONAL OBSERVATIONS OF 50 DRACONIS

The orbit of this star was published by the writer in the *Dominion Observatory Publications*, volume II, page 123, and therein it was indicated that further observations would be made to improve the value of the period which was determined solely from one season's observations.

To the 32 plates obtained in 1914 were added eight in 1915 and twelve in 1916, and from these 52 plates a solution was made which changed the value of the period from 4.120 days to 4.118 days and made small changes in the other elements. Eight plates have been obtained at odd times since then and these would indicate an even smaller value for the period, namely 4.1175, which is thus given as the best value from all the observations.

With the exception of  $T$ , the time of periastron passage, which was adjusted to conform to the revised period, the other elements from the solution of the 52 plates have been retained and are as follows:—

$$\begin{aligned}
 P &= 4.1175 \text{ days} \\
 e &= .012 \pm .009 \\
 K_1 &= 79.12 \text{ km.} \pm 0.97 \text{ km.} \\
 K_2 &= 83.90 \text{ km.} \pm 0.97 \text{ km.} \\
 \gamma &= -8.79 \text{ km.} \pm 0.49 \text{ km.} \\
 \omega_1 &= 107^\circ.6 \pm 8^\circ.9 \\
 \omega_2 &= 287^\circ.6 \pm 8^\circ.9 \\
 T &= \text{J. D. } 2,420,293.519 \pm .102 \\
 a_1 \sin i &= 4,480,000 \text{ km.} \\
 a_2 \sin i &= 4,750,600 \text{ km.} \\
 m_1 \sin^3 i &= .95 \odot \\
 m_2 \sin^3 i &= .90 \odot
 \end{aligned}$$

The probable error of a plate is  $\pm 4.6$  km. per sec. for component I, and  $\pm 4.8$  km. per sec. for component II. The table of measures following is a continuation of the table on page 123 of the volume mentioned above, but the phases there given, if used, should be revised to suit the new  $P$  and  $T$ .

ADDITIONAL MEASURES OF 50 DRACONIS

Plate	Observer*	Date	Julian Date	Phase	Component I				Component II			
					n	Wt.	Vel.	O-C	n	Wt.	Vel.	O-C
1915												
7124	C	July 26	2,420,705.770	.501	4	3	- 81.6	- 3.0	5	4	+ 58.5	- 6.7
7133	H	" 29	708.721	3.452	2	2	+ 44.0	- 0.3	4	3	- 59.7	- 5.2
7144	H	Aug. 9	719.682	2.061	5	4	- 8.5	.....	.....	.....	.....	.....
7163	Y	" 23	733.611	3.639	2	1.5	+ 45.6	+19.2	3	2	- 49.8	- 4.8
7179	Y	" 27	737.632	3.541	2	1.5	+ 24.9	-11.7	2	1.5	- 55.4	+ 1.0
7272	Y	Sept. 21	762.562	3.766	5	4	- 22.4	.....	.....	.....	.....	.....
7282	H	" 22	763.553	.639	4	3	- 83.8	+ 0.8	4	3	+ 83.2	+11.4
7290	H	" 27	768.610	1.579	2	1	- 50.9	- 8.0	2	1	+ 22.2	- 5.7
1916												
7665	H	May 24	2,421,008.625	2.779	4	3	+ 80.5	+13.3	4	2.5	-100.7	-13.4
7671	Y	" 26	010.610	.646	3	2	- 97.7	-14.3	4	2.5	+ 79.7	+ 8.6
7674	Y	" 28	012.694	2.730	2	1	+ 71.0	+ 5.6	5	3	- 84.2	+ 2.2
7682	H	June 1	016.650	2.569	8	6	+ 71.3	+14.0	8	5	- 70.0	+ 9.0
7720	H	July 6	051.595	.456	6	4	- 66.0	+ 6.3	3	2	+ 75.1	+15.7
7723	H	" 8	053.682	2.543	2	1.5	+ 59.5	+ 4.3	4	3	- 80.6	- 3.5
7726	H	" 9	054.778	3.639	1	1.	+ 27.8	- 2.9	.....	.....	.....	.....
7749	H	" 21	066.763	3.272	8	5	+ 62.8	+ 2.1	8	5	- 78.8	+ 3.7
7762	H	" 26	071.612	4.003	9	6	+ 5.9	.....	.....	.....	.....	.....
7769	H	Aug. 1	077.760	1.916	14	6.	- 4.5	.....	.....	.....	.....	.....
7785	H	" 15	091.703	3.507	5	3	+ 36.0	- 7.8	5	3	- 65.2	- 1.3
7788	H	" 16	092.573	.259	9	6	- 59.6	- 5.0	9	6	+ 49.6	+ 8.7
1917												
8210	H	June 27	407.691	2.447	2	1.5	+ 58.2	+ 5.4	3	2	- 64.2	+ 9.0
8214	H	July 2	412.661	3.300	1	1	+ 49.1	- 6.6	2	1.5	- 76.0	+ 1.0
8279	H	Sept. 3	475.603	.362	7	5	- 64.8	+ 5.1	8	5	+ 57.2	+ 0.3
8301	H	" 18	490.570	2.976	6	5	+ 73.8	+ 4.5	5	4	- 92.9	- 1.3
8307	H	" 24	496.698	.869	2	2	- 85.2	+ 2.2	2	2	+ 59.2	-15.0
1918												
8654	H	Sept. 24	861.650	3.481	3	2	+ 39.5	- 0.3	2	1.5	- 54.5	+ 5.8
8658	H	" 30	867.640	1.236	3	2	- 76.9	- 4.0	3	2	+ 45.8	-12.7
1919												
8725	H	Mar. 13	2,422,031.862	.758	1	1	-103.3	-15.5	1	1	+ 61.3	-13.5

\*C = Cannon; H = Harper; Y = Young

## MEASURES OF 50 DRACONIS (primary)

$\lambda$	7124		7133		7144		7163		7179		7272		7282			
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.		
4549					-15.3	$\frac{1}{2}$							-38.3	$\frac{1}{2}$	-111.5	$\frac{1}{2}$
4481					+5.4	$\frac{1}{2}$	+29.5	$\frac{1}{2}$					25.3	$\frac{1}{2}$		
4340			+45.9	$\frac{3}{4}$	-12.3	$\frac{1}{2}$							19.9	$\frac{1}{2}$		
4308															94.2	$\frac{1}{2}$
4250					-24.6	$\frac{1}{2}$										
4143	-89.6	$\frac{1}{2}$														
4101	62.5	$\frac{1}{2}$													79.8	$\frac{1}{2}$
4077	86.8	$\frac{1}{2}$							+18.2	$\frac{1}{2}$			29.7	$\frac{3}{4}$		
4063	79.8	$\frac{1}{2}$							+21.0	$\frac{3}{4}$						
4045	101.5	$\frac{1}{2}$														
3933	-82.8	$1\frac{1}{2}$	+36.0	1	-15.8	$\frac{1}{2}$	+46.8	$\frac{1}{2}$					-21.9	$\frac{3}{4}$	-80.9	$1\frac{1}{2}$
Weighted mean	-85.07		+40.30		-12.52		+41.03		+20.30				-26.80		-88.20	
$V_a$	+3.87		+4.02		+4.38		+4.83		+4.93				+4.72		+4.69	
$V_s$	-.09		-.05		-.05		-.03		-.05				-.03		-.05	
Curv.	-.29		-.28		-.28		-.28		-.28				-.28		-.28	
Radial Velocity	-81.6		+44.0		-8.5		+45.6		+24.9				-22.4		-83.8	

MEASURES OF 50 DRACONIS (primary)—Continued

λ	7290		7665		7671		7674		7682		7720		7723	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4549									+85.5	½	-72.7	½	+64.8	½
4520									68.1	½				
4481			+71.1	½	-111.2	½	+78.6	½			62.8	½	+49.1	½
4340									49.3	½	50.3	¼		
4271			116.1	¼							79.7	½		
4260									61.6	¼				
4227					79.6	¼								
4143									70.3	¼				
4101							+64.6	¼						
4077									83.1	¼				
4045	-61.8	¼	88.9	¼	-86.2	¼			71.5	½	63.5	¼		
3933	-48.3	¼	+70.7	½					+79.5	¼	-84.5	¼		
Weighted mean	- 55.05		+ 81.33		- 97.00		+ 71.60		+ 71.57		- 68.50		+ 56.95	
V <sub>s</sub>	+ 4.50		- 0.70		- 0.53		- 0.36		- 0.03		+ 2.70		+ 2.84	
V <sub>d</sub>	- .06		+ .12		+ .08		+ .06		+ .06		+ .05		+ .02	
Curv.	- .28		- .28		- .28		- .28		- .28		- .28		- .28	
Radial Velocity	- 50.9		+ 80.5		- 97.7		+ 71.0		+ 71.3		- 66.0		+ 59.5	

## MEASURES OF 50 DRACONIS (primary)—Continued

$\lambda$	7726		7749		7762		7769		7785		7788		8210	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4584					-12.8	$\frac{1}{2}$								
4563			+64.7	$\frac{1}{2}$			+0.5	$\frac{1}{2}$						
4549					+17.1	$\frac{1}{2}$	-24.6	$\frac{2}{3}$						
4481			76.1	$\frac{1}{2}$	+2.9	$\frac{1}{2}$	-19.4	$\frac{2}{3}$			-58.8	$\frac{1}{2}$	+58.0	$\frac{1}{2}$
4468			48.7	$\frac{1}{2}$			+1.2	$\frac{1}{2}$						
4415											56.9	$\frac{1}{2}$		
4352							-9.5	$\frac{2}{3}$						
4340			55.4	$\frac{1}{2}$			-12.5	$\frac{1}{2}$			64.9	$\frac{1}{2}$		
4325							-12.0	$\frac{2}{3}$			77.4	$\frac{1}{2}$		
4308					+4.0	$\frac{1}{2}$	+1.2	$\frac{1}{2}$						
4294							-11.7	$\frac{2}{3}$						
4289							+8.6	$\frac{1}{2}$						
4271					-9.2	$\frac{1}{2}$					54.9	$\frac{1}{2}$		
4236					+8.6	$\frac{1}{2}$	-21.1	$\frac{1}{2}$						
4233			73.2	$\frac{1}{2}$	+20.9	$\frac{1}{2}$	-1.0	$\frac{2}{3}$	+29.4	$\frac{1}{2}$	47.4	$\frac{1}{2}$		
4215					-16.6	$\frac{1}{2}$								
4143			63.5	$\frac{1}{2}$										
4077					+11.8	$\frac{1}{2}$	-10.7	$\frac{2}{3}$						
4063							+1.0	$\frac{1}{2}$	22.7	$\frac{1}{2}$	67.7	$\frac{1}{2}$		
4045			45.2	$\frac{1}{2}$					21.8	$\frac{1}{2}$	71.9	$\frac{1}{2}$		
4005									52.2	$\frac{1}{2}$				
3933	+25.2	$\frac{1}{2}$	+61.2	$\frac{1}{2}$					+32.3	$\frac{1}{2}$	-80.0	$\frac{1}{2}$	+53.3	$\frac{1}{2}$
Weighted mean	+ 25.20		+ 59.44		+ 2.22		- 8.48		+ 31.68		- 64.03		+ 56.42	
$V_a$	+ 2.91		+ 3.65		+ 3.90		+ 4.19		+ 4.68		+ 4.71		+ 2.04	
$V_d$	- .04		- .05		+ .03		- .05		- .05		+ .02		+ .03	
Curv.	- .28		- .28		- .28		- .28		- .28		- .28		- .28	
Radial Velocity	+ 27.8		+ 62.8		+ 5.9		- 4.5		+ 36.0		- 59.6		+ 58.2	

MEASURES OF 50 DRACONIS (primary)—Continued

λ	8214		8279		8301		8307		8654		8658		8725	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4549			-76.8	$\frac{1}{2}$			-122.3	$\frac{1}{2}$			-104.2	$\frac{1}{2}$		
4481	+46.7	$\frac{1}{4}$	66.4	$\frac{1}{2}$	+64.9	1	-67.6	$\frac{3}{4}$	+31.2	$\frac{1}{4}$	72.1	$\frac{1}{4}$	-98.2	$\frac{1}{2}$
4415					78.6	$\frac{1}{2}$								
4340			68.8	$\frac{1}{4}$	70.9	1			12.5	$\frac{1}{2}$				
4143					64.0	$\frac{1}{2}$								
4101			74.2	$\frac{1}{2}$	73.2	$\frac{1}{2}$								
4077			58.1	$\frac{1}{2}$										
4045			69.2	$\frac{1}{4}$										
3933			-75.2	$\frac{1}{4}$	+16.3	$\frac{1}{2}$			+59.2	$\frac{1}{2}$	-78.2	$\frac{1}{4}$		
Weighted mean	+ 46.70		- 69.45		+ 69.32		- 89.44		+ 34.88		- 80.96		- 98.20	
V <sub>0</sub>	+ 2.40		+ 4.93		+ 4.78		+ 4.62		+ 4.84		+ 4.43		- 4.91	
V <sub>a</sub>	+ .03		- .02		- .04		- .08		+ .07		- .07		+ .09	
Curv.	- .28		- .28		- .28		- .28		- .28		- .28		- .28	
Radial Velocity	+ 49.1		- 64.8		+ 73.8		- 85.2		+ 39.5		- 76.9		- 103.3	

## MEASURES OF 50 DRACONIS (secondary)—Continued

$\lambda$	7124		7124		7133		7163		7179		7282		7290	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4549											+70.0	$\frac{1}{2}$		
4534							-51.2	$\frac{1}{2}$						
4340					-59.3	$\frac{1}{2}$								
4308											82.4	$\frac{1}{2}$		
4271					72.3	$\frac{1}{2}$								
4143			+65.9	$\frac{1}{2}$			60.3	$\frac{1}{2}$						
4101	+67.5	$\frac{1}{2}$									84.3	$\frac{1}{2}$		
4077			63.9	$\frac{1}{2}$					-69.0	$\frac{1}{2}$				
4063	50.8	$\frac{1}{2}$	44.3	$\frac{1}{2}$										
4045	44.4	1	47.0	$\frac{1}{2}$					-57.0	$\frac{1}{2}$			+24.3	$\frac{1}{2}$
3968					67.9	$\frac{1}{2}$								
3933	+59.0	$1\frac{1}{2}$	+58.3	$1\frac{1}{2}$	-61.8	1	-54.5	$\frac{1}{2}$			+78.8	$1\frac{1}{2}$	+11.6	$\frac{1}{2}$
Weighted mean	+ 53.44		+ 56.54		- 63.36		- 54.35		- 60.00		+ 78.84		+ 18.00	
$V_s$	+ 3.87		+ 3.87		+ 4.02		+ 4.83		+ 4.93		+ 4.69		+ 4.50	
$V_d$	- .09		- .09		- .05		- .03		- .05		- .05		- .06	
Curv.	- .28		- .28		- .28		- .28		- .28		- .28		- .28	
Radial Velocity	+ 57.0		+ 60.0		- 59.7		- 49.8		- 55.4		+ 83.2		+ 22.2	

MEASURES OF 50 DRACONIS (secondary)—Continued

$\lambda$	7665		7671		7674		7682		7720		7723		7726	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4549							-64.4	$\frac{1}{2}$	+51.7	$\frac{1}{2}$	-93.2	$\frac{1}{2}$		
4520							62.9	$\frac{1}{2}$						
4481	-117.0	$\frac{1}{2}$	+35.0	$\frac{1}{2}$	-78.6	$\frac{1}{2}$			74.9	$\frac{1}{2}$	84.6	$\frac{1}{2}$		
4340							60.5	$\frac{1}{2}$						
4325											71.8	$\frac{1}{2}$		
4308	84.2	$\frac{1}{2}$			91.3	$\frac{1}{2}$								
4271					106.0	$\frac{1}{2}$								
4260					80.2	$\frac{1}{2}$	81.8	$\frac{1}{2}$						
4233			95.0	$\frac{1}{2}$										
4227			73.6	$\frac{1}{2}$										
4143							90.3	$\frac{1}{2}$						
4077							83.5	$\frac{1}{2}$						
4045	103.8	$\frac{1}{2}$	+63.3	$\frac{1}{2}$			51.9	$\frac{1}{2}$			-83.0	$\frac{1}{2}$		
3933	-94.3	$\frac{1}{2}$			-54.7	$\frac{1}{2}$	-77.8	$\frac{1}{2}$	+91.2	$\frac{1}{2}$			-39.2	$\frac{1}{2}$
Weighted mean	-99.80		+80.40		-83.60		-69.73		+72.60		-83.15		-39.20	
$V_s$	-0.70		-0.53		-0.36		-0.03		+2.70		+2.84		+2.91	
$V_d$	+0.12		+0.08		+0.06		+0.06		+0.05		+0.02		-0.04	
Curv.	-0.28		-0.28		-0.28		-0.28		-0.28		-0.28		-0.28	
Radial Velocity	-100.7		+79.7		-84.2		-70.0		+75.1		-80.6		-36.6	

## MEASURES OF 50 DRACONIS (secondary)—Continued

$\lambda$	7749		7785		7788		8210		8214		8279		8301	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4563	-98.0	$\frac{1}{2}$												
4549											+40.9	$\frac{1}{2}$		
4481	74.8	$\frac{1}{2}$			+28.1	$\frac{1}{2}$	-74.5	$\frac{1}{2}$	-82.1	$\frac{1}{2}$	34.2	$\frac{1}{2}$	-105.8	$\frac{1}{2}$
4468	101.4	$\frac{1}{2}$												
4415					38.2	$\frac{1}{2}$								
4352	64.0	$\frac{1}{2}$												
4340	84.1	$\frac{1}{2}$			42.6	$\frac{1}{2}$					38.4	$\frac{1}{2}$	93.8	1
4325					68.1	$\frac{1}{2}$			-75.5	$\frac{1}{2}$				
4271					54.3	$\frac{1}{2}$								
4233	58.8	$\frac{1}{2}$	-66.8	$\frac{1}{2}$	54.1	$\frac{1}{2}$								
4143	67.8	$\frac{1}{2}$											117.8	$\frac{1}{2}$
4101											75.0	$\frac{1}{2}$	85.7	$\frac{1}{2}$
4077							56.6	$\frac{1}{2}$			53.5	$\frac{1}{2}$		
4063			54.9	$\frac{1}{2}$	59.4	$\frac{1}{2}$					63.1	$\frac{1}{2}$		
4045			70.4	$\frac{1}{2}$	40.3	$\frac{1}{2}$					40.1	$\frac{1}{2}$		
4005			74.4	$\frac{1}{2}$										
3933	-87.7	$\frac{1}{2}$	-81.0	$\frac{1}{2}$	+30.0	$\frac{1}{2}$	-76.5	$\frac{1}{2}$			+60.0	$\frac{1}{2}$	-97.6	$\frac{1}{2}$
Weighted mean	-82.16		-69.50		+45.10		-66.00		-78.10		+52.53		-97.40	
$V_s$	+3.65		+4.68		+4.71		+2.04		+2.40		+4.93		+4.78	
$V_d$	-.05		-.05		+.02		+.03		+.03		-.02		-.04	
Curv.	-.28		-.28		-.28		-.28		-.28		-.28		-.28	
Radial Velocity	-78.8		-65.2		+49.6		-64.2		-76.0		+57.2		-92.9	



## BOSS 5070

(1900,  $\alpha = 19^{\text{h}} 47^{\text{m}} \cdot 2$ ,  $\delta = +40^{\circ} 20'$ , mag. 5.62, type A)

This star was announced as a spectroscopic binary by Adams in the *Publications of the Astronomical Society of the Pacific*, vol. 26, page 261. From the measures, the period is evidently short. The *K* line of calcium does not share in the large oscillations of the other lines. Sufficient plates for an orbit have been secured at Allegheny and the star is dropped from our list.

Our measures follow.

Plate	Date, G.M.T.	Velocity	<i>n</i>	Weight
6804.....	1915, Feb. 18.941	-23	9	3
6842.....	Mar. 3.910	+19	10	4
6928.....	April 14.803	+88	7	3
8285.....	1917, Sept. 6.747	-86	5	2
8290.....	" 8.527	+4	6	2
8294.....	" 11.631	+67	7	3
8297.....	" 12.653	+73	4	2

## MEASURES OF BOSS 5070

$\lambda$	6804		6842		6928		8285		8290		8294		8297	
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4567.967	+11.5	$\frac{1}{2}$	+ 6.6	$\frac{1}{2}$	+ 99.5	$\frac{1}{2}$	.....	.....	.....	.....	.....	.....	.....	.....
4552.762	-26.3	$\frac{1}{2}$	+24.4	$\frac{1}{2}$	+101.6	$\frac{1}{2}$	.....	.....	+40.9	$\frac{1}{2}$	+77.4	$\frac{1}{2}$	+ 58.4	$\frac{1}{2}$
4481.400	-12.0	$\frac{1}{2}$	.....	.....	.....	.....	.....	.....	.....	.....	.....	.....	.....	.....
4471.676	-33.6	$\frac{2}{3}$	+18.7	$\frac{1}{2}$	+ 69.6	$\frac{1}{2}$	-73.1	$\frac{2}{3}$	-11.0	$\frac{1}{2}$	+82.1	$\frac{2}{3}$	+115.5	$\frac{2}{3}$
4388.100	.....	.....	+28.5	$\frac{1}{2}$	+ 75.4	$\frac{1}{2}$	-97.0	$\frac{1}{2}$	+18.3	$\frac{1}{2}$	+65.3	$\frac{1}{2}$	+ 70.0	$\frac{1}{2}$
4340.634	-54.4	$\frac{1}{2}$	- 4.5	$\frac{1}{2}$	+ 43.0	$\frac{1}{2}$	-66.4	$\frac{1}{2}$	+16.1	$\frac{1}{2}$	+77.7	$\frac{2}{3}$	+ 58.9	$\frac{1}{2}$
4325.939	.....	.....	+21.7	$\frac{1}{2}$	.....	.....	.....	.....	.....	.....	.....	.....	.....	.....
4267.301	-23.8	$\frac{1}{2}$	+22.3	$\frac{1}{2}$	.....	.....	.....	.....	.....	.....	+94.8	$\frac{1}{2}$	.....	.....
4143.928	-16.1	$\frac{1}{2}$	- 1.9	$\frac{1}{2}$	.....	.....	-78.9	$\frac{1}{2}$	.....	.....	+52.0	$\frac{1}{2}$	.....	.....
4101.890	-19.0	$\frac{1}{2}$	+11.6	$\frac{1}{2}$	+ 54.6	$\frac{1}{2}$	-72.7	$\frac{1}{2}$	- 1.9	$\frac{1}{2}$	.....	.....	.....	.....
4026.352	.....	.....	.....	.....	+ 88.4	$\frac{1}{2}$	.....	.....	+ 1.5	$\frac{1}{2}$	+72.0	$\frac{1}{2}$	.....	.....
3933.825	[ - 5.8	$\frac{1}{2}$	[ -29.0	$\frac{1}{2}$	[ - 6.5	$\frac{1}{2}$	.....	.....	.....	.....	.....	.....	.....	.....
Weighted mean	- 26.87		+ 12.31		+ 74.47		- 77.56		+ 12.76		+ 76.56		+ 82.80	
<i>V</i> <sub>0</sub>	+ 4.07		+ 7.20		+ 13.99		- 7.74		- 8.13		- 8.78		- 8.99	
<i>V</i> <sub>d</sub>	+ .22		+ .21		+ .21		- .21		+ .07		- .11		- .13	
Curv.	- .28		- .28		- .28		- .28		- .28		- .28		- .28	
Radial Velocity	- 22.9		+ 19.4		+ 88.4		- 85.8		+ 4.4		+ 67.4		+ 73.4	

13 VULPECULÆ

(1900,  $\alpha = 19^h 49^m \cdot 2$ ,  $\delta = +23^\circ 50'$ , mag. 4.50, type A)

This star was announced a binary by Lee, in *Astrophysical Journal*, XXXII, 307, from eight plates giving a range from - 15 to - 36. Campbell used as the velocity for the star, - 28. While the lines are in general fairly sharp there is a bare suspicion that  $H\gamma$  on plate 8221 is double.

Plate	Date, G.M.T.	Velocity	Lines	Weight
8202.....	1917, June 18.714	-37.8	1	1
8218.....	July 5.740	-24.2	4	2
8221.....	" 6.690	-30.8	2	1
8222.....	" 14.678	-30.4	5	2
8225.....	" 15.777	-32.5	5	2
8228.....	" 16.691	-24.1	1	1

MEASURES OF 13 VULPECULÆ

$\lambda$	8202		8218		8221		8222		8225		8228			
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4584.018							-18.6	$\frac{1}{2}$						
4549.743			-15.6	$\frac{1}{4}$			+35.3	$\frac{1}{4}$	-42.1	$\frac{1}{2}$				
4481.477	-50.9	$\frac{3}{4}$	44.0	$\frac{1}{2}$	-43.1	$\frac{1}{2}$	+51.5	$\frac{1}{2}$	+24.1	$\frac{1}{2}$	-28.1	$\frac{3}{4}$		
4340.645			25.8	$\frac{1}{4}$			+33.9	$\frac{1}{2}$	+40.0	$\frac{1}{4}$				
4233.425									+41.5	$\frac{1}{4}$				
3933.825			-31.2	$\frac{1}{2}$	-28.6	$\frac{1}{4}$	-36.6	$\frac{1}{2}$	-43.2	$\frac{1}{4}$				
Weighted mean	- 50.90		- 31.93		- 38.27		- 35.15		- 36.70		- 28.10			
$V_a$	+ 13.20		+ 7.99		+ 7.64		+ 4.96		+ 4.58		+ 4.20			
$V_d$	+ .14		+ .02		+ .08		+ .08		+ .08		+ .04			
Curv.	- .28		- .28		- .28		- .28		- .28		- .28			
Radial Velocity	- 37.8		- 24.2		- 30.8		- 30.4		- 32.5		- 24.1			

## BOSS 5535

(1900,  $\alpha = 21^{\text{h}} 28^{\text{m}} \cdot 3$ ,  $\delta = + 60^{\circ} 01'$ , mag. 5.52, type A)

This star is in Kapteyn's Area No. 18. It is listed as of A-type in *Harvard Annals*, vol. 50, but would more properly fall under the classification B2. Besides the hydrogen and helium series, there are the characteristic lines of this type at  $\lambda\lambda$  4089 and 4649 variously ascribed to argon, silicon and other substances. The third member of the group, that at  $\lambda$  4116, does not appear on the plates. The three silicon lines  $\lambda\lambda$  4575.52, 4568.13 and 4552.89 are seen in the spectrum, the latter two being 0.5 as intense as  $H\gamma$  or the helium  $\lambda$  4471. The wave-lengths indicated for them are somewhat greater than the generally accepted values, but were adjusted to agree with the velocities obtained from the hydrogen and helium series. Other absorption lines, whose normal wave-lengths are approximately  $\lambda\lambda$  4366.9, 4620.5, 4630.7 and 4641.4, were noted but not used in the results.

Plate	Date, G.M.T.	Number of Lines	Velocity
8394.....	1917, Dec. 28.465	5	-28.3
8404.....	1918, Jan. 2.481	5	-18.7
8410.....	" 4.455	6	-19.4
8672.....	Oct. 23.605	7	-33.8

## MEASURES OF BOSS 5535

$\lambda$	8394		8404		8410		8672							
	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.	Vel.	Wt.
4568.13-	-14.0	$\frac{1}{2}$	+19.1	$\frac{1}{2}$	-9.7	$\frac{1}{2}$	-38.6	$\frac{1}{2}$						
4552.89-	12.3	$\frac{1}{2}$	-1.4	$\frac{1}{2}$	-5.3	$\frac{1}{2}$	31.0	$\frac{1}{2}$						
4471.676	19.7	$\frac{2}{3}$	-29.0	$\frac{1}{2}$	+1.8	$\frac{1}{2}$	22.3	$\frac{2}{3}$						
4388.100	19.6	$\frac{1}{2}$	-2.1	$\frac{1}{2}$	-0.9	$\frac{1}{2}$	26.4	$\frac{2}{3}$						
4340.634	-11.2	$\frac{1}{2}$	+4.6	$\frac{1}{2}$	-16.7	$\frac{1}{2}$	34.2	$\frac{2}{3}$						
4143.928					-18.4	$\frac{1}{2}$	17.3	$\frac{1}{2}$						
4121.016							-35.5	$\frac{1}{2}$						
Weighted mean	-15.84		-6.20		-6.96		-29.01							
$V_s$	-12.07		-12.07		-12.04		-4.42							
$V_d$	-11		-14		-12		-09							
Curv.	-28		-28		-28		-28							
Radial Velocity	-28.3		-18.7		-19.4		-33.8							

Dominion Observatory  
Ottawa  
April, 1919.

